

# ISS Plasma Environment: Status of CCMC Products for ISS Mission Ops

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**GSFC Space Weather and NASA Robotic Missions Workshop**  
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# ISS Space Weather Needs\*

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- Solar activity/thermosphere density prediction and satellite torque/drag predictions for:
  - Mission planning and controllability/real-time operations
  - MM/OD environment evolution
- MSFC meteor storm severity predictions driven by the Perseids on Mir events of 12 August 1993
- Role of solar/geomagnetic activity/thermosphere in managing ISS crew ionizing radiation dose exposure
- ISS interaction with auroral particle precipitation – why this isn't a problem or do we need forecasting?
- Monitor for changes in the SAA altitude structure and geographic extent for crew IR dose management
- Ionospheric Ne, Te values along ISS orbit for characterizing ISS charging hazards:
  - Near real time Ne, Te data
  - Well validated real time model Ne, Te output

\*from S. Koontz/ISS Environments Manager



# Introduction

## Overview

- Background
  - ISS interaction with plasma environment
  - Charging hazards to vehicle and crew
  - ISS Program hazard mitigation strategies
- ISS plasma environment monitoring
  - Floating Potential Measurement Unit (FPMU)
- CCMC CTIPe real time plasma model
  - Examples of success
  - Issues in work
- Summary and future needs

ISS/FPMU:	JSC	Steve Koontz (ISS Environments Lead), John Alred, Jack Bacon
	MSFC	Mike Chandler, Victoria Coffey, Joseph Minow, Paul Craven, Todd Schneider, Ken Wright (UAH)
CTIPe model:	Boeing	Drew Hartman, Leonard Kramer, Ron Mikatarian, Danny Schmidt
	NOAA	Tim Fuller-Rowell
CCMC team:	GSFC	Ja Soon Shim (UMBC/NASA), Lutz Rastaetter (NASA), Masha Kuznetsova (NASA)



# ISS Structure Potential Variations and Hazards

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- ISS potential varies in low Earth orbit environment due to:

## Current collection

- Current collection from ambient plasma
- 160 V US solar array
- Visiting vehicle (high voltage) solar arrays
- Operation of payloads that emit current sources

## Observed Voltage Ranges to Date

- 0.1 to -0.5 volts
- 20 to -90 volts
- 10 volts
- +10 to +25 volts

## Inductive potentials

- $(v \times B) \cdot L$  due to motion across geomagnetic field
- $E \cdot L$  due to ionospheric electric fields
- Auroral electrons

+/-40 volts

few volts

-20 volts

- Hazards to vehicle and crew
  - ISS-EVA-305: long term degradation of thin dielectric surface thermal control coatings due to arcing ...EVA touch temp violations (eventually)
    - Hazard marginalized by test and analysis - no controls needed
  - ISS-EVA-312: EVA electric shock
    - Hazard 1 - Catastrophic at floating potentials more negative than -40V
    - Hazard 2 - Critical to catastrophic at positive floating potentials ( $> 0V$ )
    - Hazard 3 – critical to catastrophic ISS electrical power short through EVA crew to ground
      - Plasma is a secondary cause – one circuit closure pathway

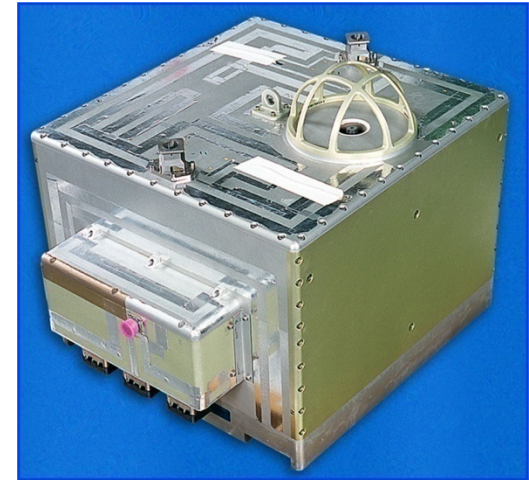




# ISS Plasma Hazard Management

ISS Program controls plasma hazards through a process of active potential control, operational mitigation strategies, environment monitoring and characterization, and probabilistic risk assessment

- **Plasma Contactor Units (PCUs)**
  - Provides active ISS “ground” by dissipating surface charges to space
  - Two redundant PCU units provides single fault failure tolerance, two required for EVA
- **Operational control using ISS flight attitude, solar array wing angle, and solar array shunt state**
  - Manages solar array and magnetic induction charging
  - Provides two fault tolerance
- **Floating Potential Measurement Unit (FPMU)**
  - Provides validated measurements of ISS floating potential and ionospheric Ne, Te along ISS orbit
- **Plasma Interaction Model (PIM)**
  - ISS charging model validated with FPMU data
  - Predicts charging hazard severity and frequency of occurrence





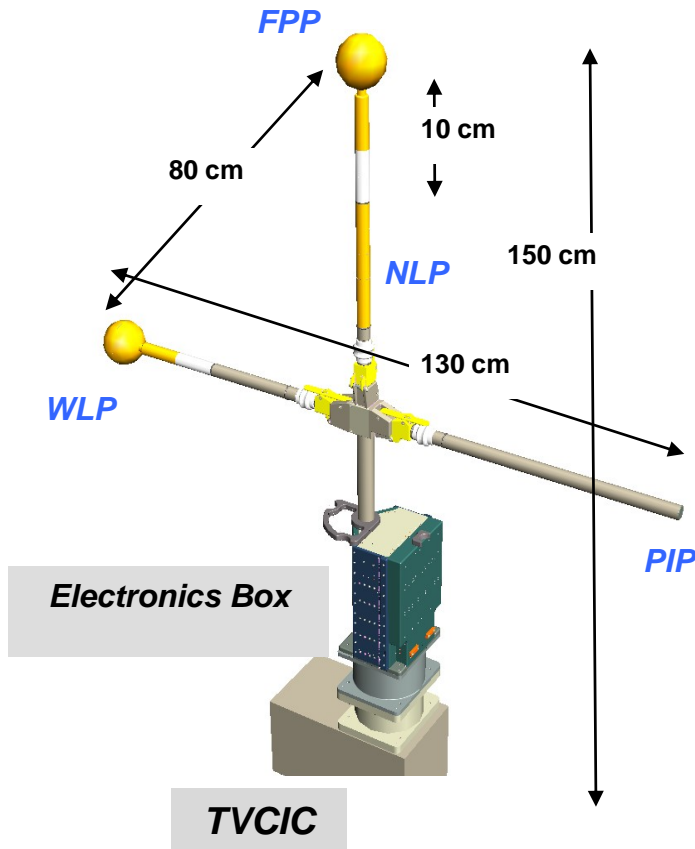
# Floating Potential Measurement Unit (FPMU)

**FPP: Floating Potential Probe**

**WLP: Wide-sweep Langmuir Probe**

**NLP: Narrow-sweep Langmuir Probe**

**PIP: Plasma Impedance Probe**



## Role:

- **Validation of PIM**
- **Assess PV array variability**
- **Interpreting IRI predictions**
- **Characterize ISS charging**

Sensor	Measured Parameter	Rate (Hz)	Effective Range
FPP	$V_F$	128	-180 V to +180 V
WLP	$N$ $T_e$ $V_F$	1	$10^9 \text{ m}^{-3}$ to $5 \cdot 10^{12} \text{ m}^{-3}$ 500 K to ~10000 K -20 V to 80 V
NLP	$N$ $T_e$ $V_F$	1	$10^9 \text{ m}^{-3}$ to $5 \cdot 10^{12} \text{ m}^{-3}$ 500 K to ~10000 K -180V to +180 V
PIP	$N$	512	$1.1 \cdot 10^{10} \text{ m}^{-3}$ to $4 \cdot 10^{12} \text{ m}^{-3}$

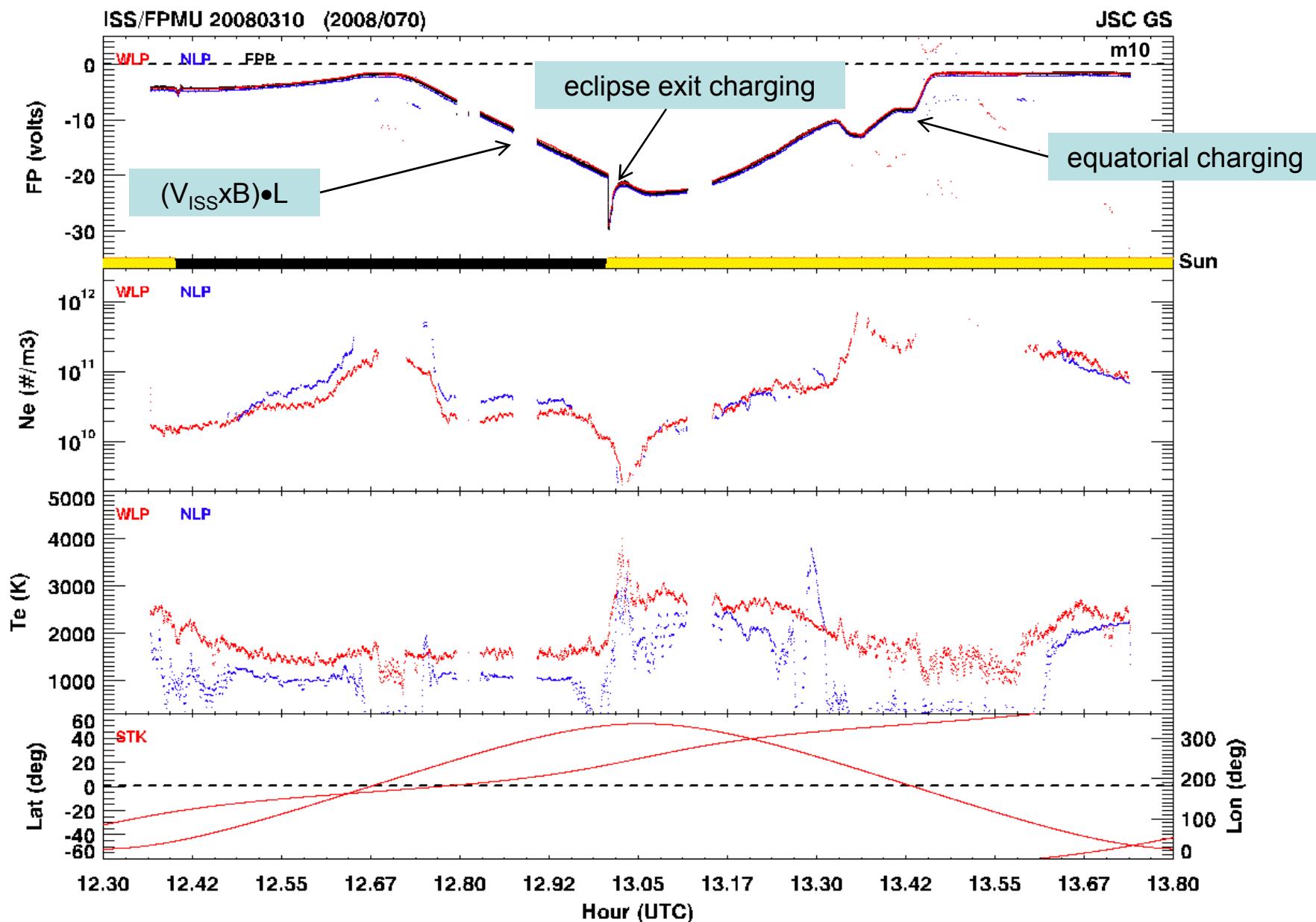
[Wright et al., 2008; Barjatya et al., 2009]







# Characterizing ISS Environments, Charging







# Alternative Ne, Te Data and/or Model Sources

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We have in-situ FPMU Ne, Te measurements along ISS orbit...

...but we are interested in identifying independent Ne, Te data sources from both measurements and models appropriate for ISS altitudes

- FPMU validation against independent measurements, models
  - Incoherent scatter radar Ne, Te, ionosonde Ne, TIMED/GUVI Ne
  - IRI, GAIM model Ne, Te
- FPMU data unavailable during EVA, docking, and other operations with higher Ku band video downlink priority
  - Real time data, models may be useful to provide coverage during these periods
- FPMU operated on campaign basis (~25 to 30% of year)
  - Well validated models or alternate data sources can provide environment characterization data between FPMU runs
- Contingency planning in case of FPMU failure
  - Default to current “worst case” analysis for EVA planning...but that impacts ISS power availability
  - Alternative data, validated models could provide operations relief to power constraint

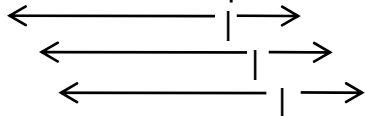


# CCMC Real-time Ionosphere Ne, Te for ISS

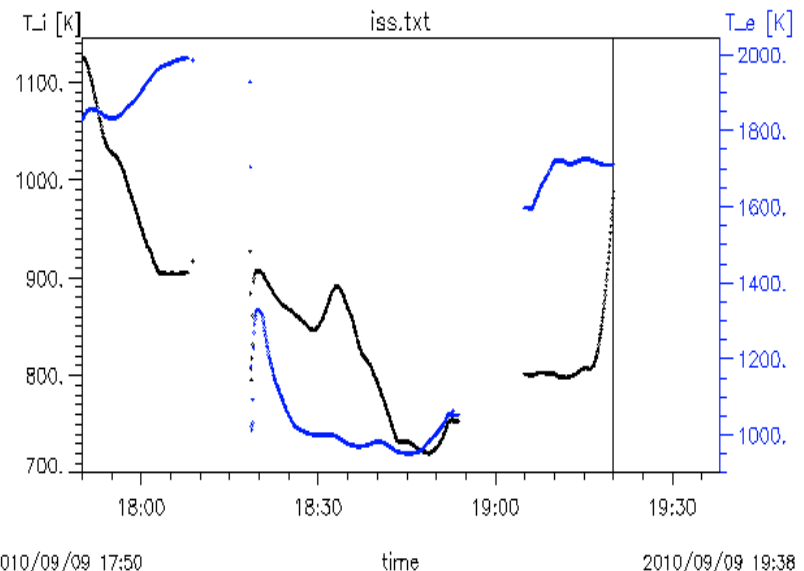
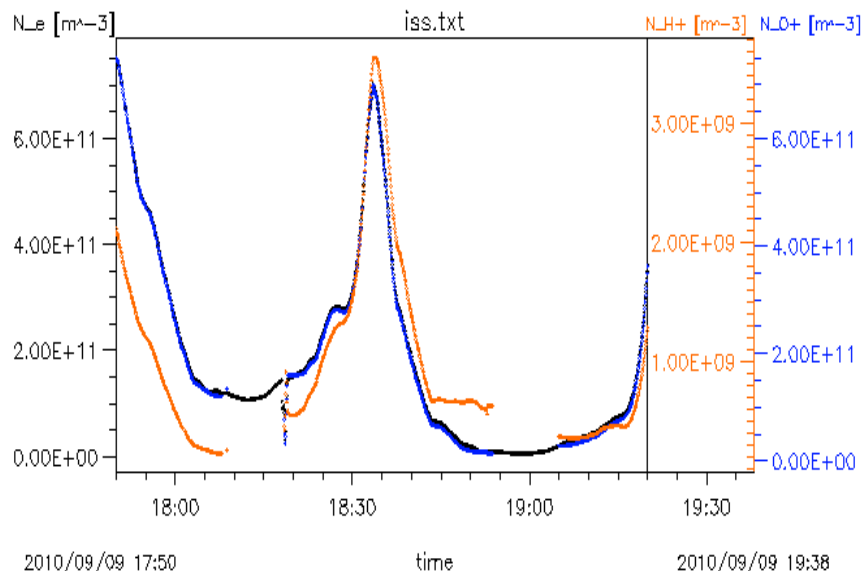
## Coupled Thermosphere Ionosphere Plasmasphere Electrodynamics (CTIPe) Model

- CCMC implemented real time CTIPe model in spring 2010 (CTIPe\_RT) with output specific for ISS orbit
- ISS ephemeris from GSFC/SSCWeb
- New record every 10 minutes gives 90 minutes of data at 5 sec time steps

-70 min from file epoch to +20 min



```
CTIPe_at_ISS_20100909_192000.txt                                09/09/2010 07:01PM
# Data printout from CCMC-simulation: version 1.1
# Data type: CTIP ionosphere/thermosphere
# Run name: 2010-09 Missing data: -1.100E+12
# Coordinate System: GEO
# fixed dipole tilt angles used: SM-GSM: 0.00000 GSM-GSE: 0.00000
# Satellite Track: iss
# Output data: field with 1x1081=1081 elements
#YYYYMM DD HH MM Sec lon lat IP N_e N_O+ N_H+ T_i T_e
# year month day h m s [deg] [deg] [km] [m^-3] [m^-3] [m^-3] [K] [K]
2010 09 09 17 50 0.000 254.4 -9.250 351.5 7.522E+11 7.501E+11 2.108E+09 1125. 1828.
2010 09 09 17 50 5.000 254.6 -8.994 351.5 7.494E+11 7.473E+11 2.089E+09 1125. 1831.
2010 09 09 17 50 10.000 254.8 -8.738 351.4 7.465E+11 7.444E+11 2.069E+09 1125. 1834.
2010 09 09 17 50 15.000 254.9 -8.483 351.4 7.434E+11 7.414E+11 2.050E+09 1125. 1837.
2010 09 09 17 50 20.000 255.1 -8.227 351.3 7.402E+11 7.382E+11 2.030E+09 1124. 1840.
2010 09 09 17 50 25.000 255.3 -7.971 351.3 7.366E+11 7.346E+11 2.010E+09 1124. 1843.
2010 09 09 17 50 30.000 255.5 -7.715 351.2 7.312E+11 7.292E+11 1.989E+09 1123. 1844.
2010 09 09 17 50 35.000 255.7 -7.459 351.1 7.259E+11 7.239E+11 1.968E+09 1122. 1846.
2010 09 09 17 50 40.000 255.9 -7.203 351.1 7.205E+11 7.186E+11 1.947E+09 1120. 1848.
2010 09 09 17 50 45.000 256.1 -6.947 351.0 7.151E+11 7.132E+11 1.927E+09 1119. 1850.
----- (records deleted) -----
2010 09 09 19 20 0.000 227.5 -14.02 352.8 3.634E+11 3.621E+11 1.289E+09 989.1 1710.
```





# Example CTIPe\_RT Daily Output

## CTIPe\_RT output at CCMC:

Integrated Space Weather Analysis System (iSWA)

<http://iswa.gsfc.nasa.gov/iswa/iSWA.html>

Anonymous ftp

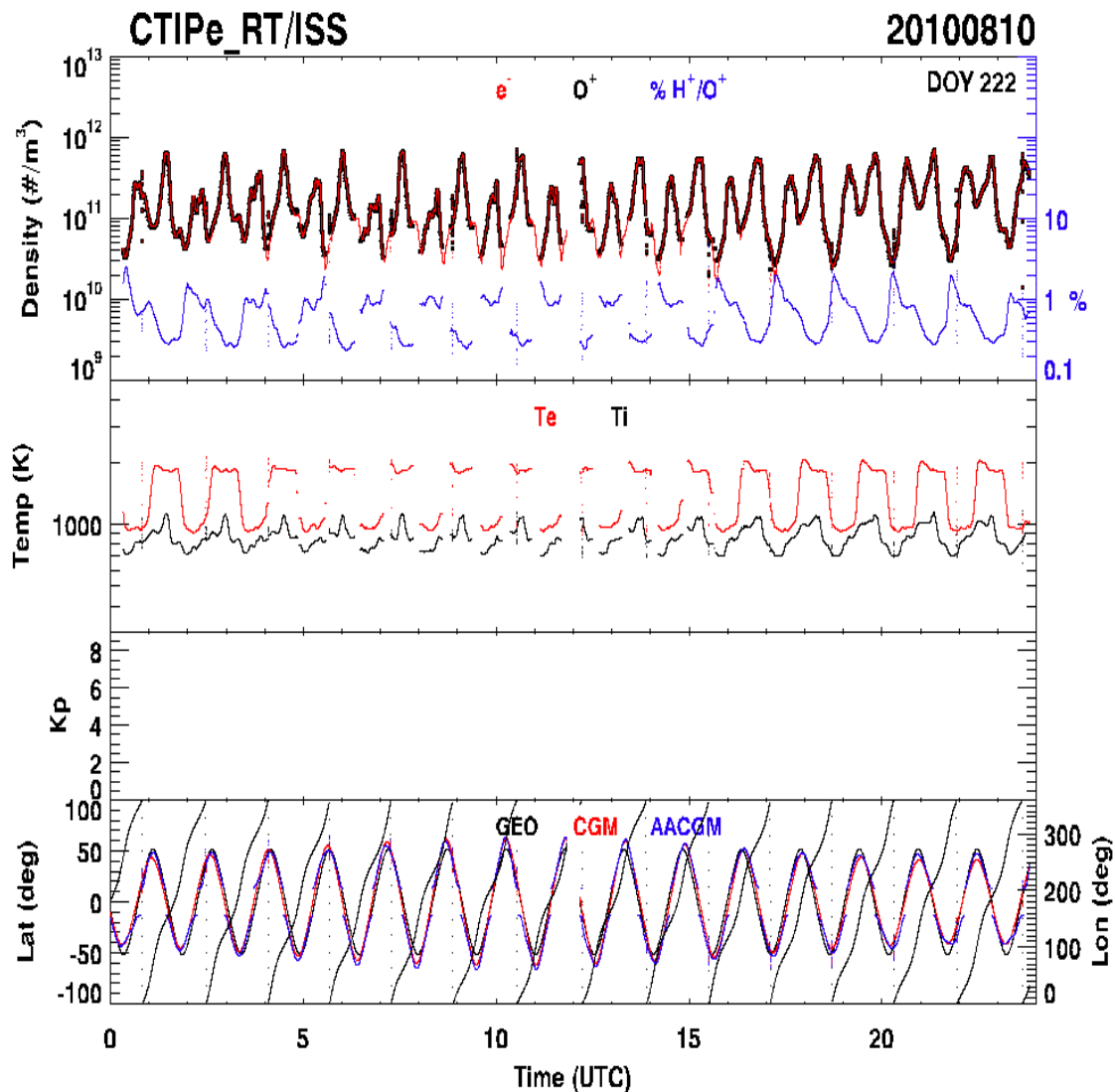
<ftp://hanna.ccmc.gsfc.nasa.gov/>

CTIPe Model Description:

<http://ccmc.gsfc.nasa.gov/models/modelinfo.php?model=CTIPe>

## MSFC is evaluating CTIPe\_RT for possible ISS ops use:

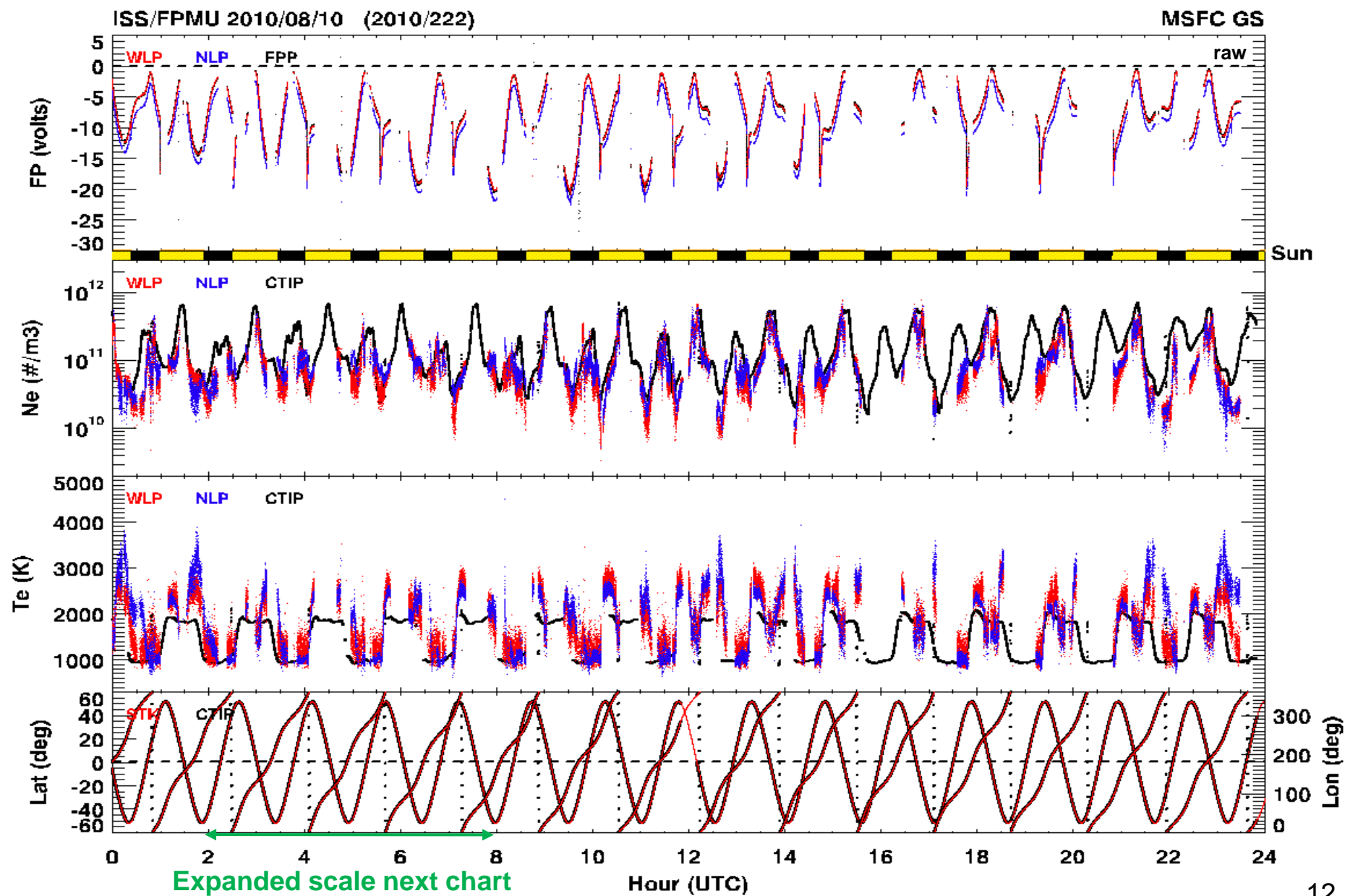
- Periodically download text output files and process into daily data sets retaining the unique records
- Compare CTIPe\_RT Ne, Te with measurements from FPMU
- This is a work in progress, both for ISS and CCMC! Only preliminary results shown here...





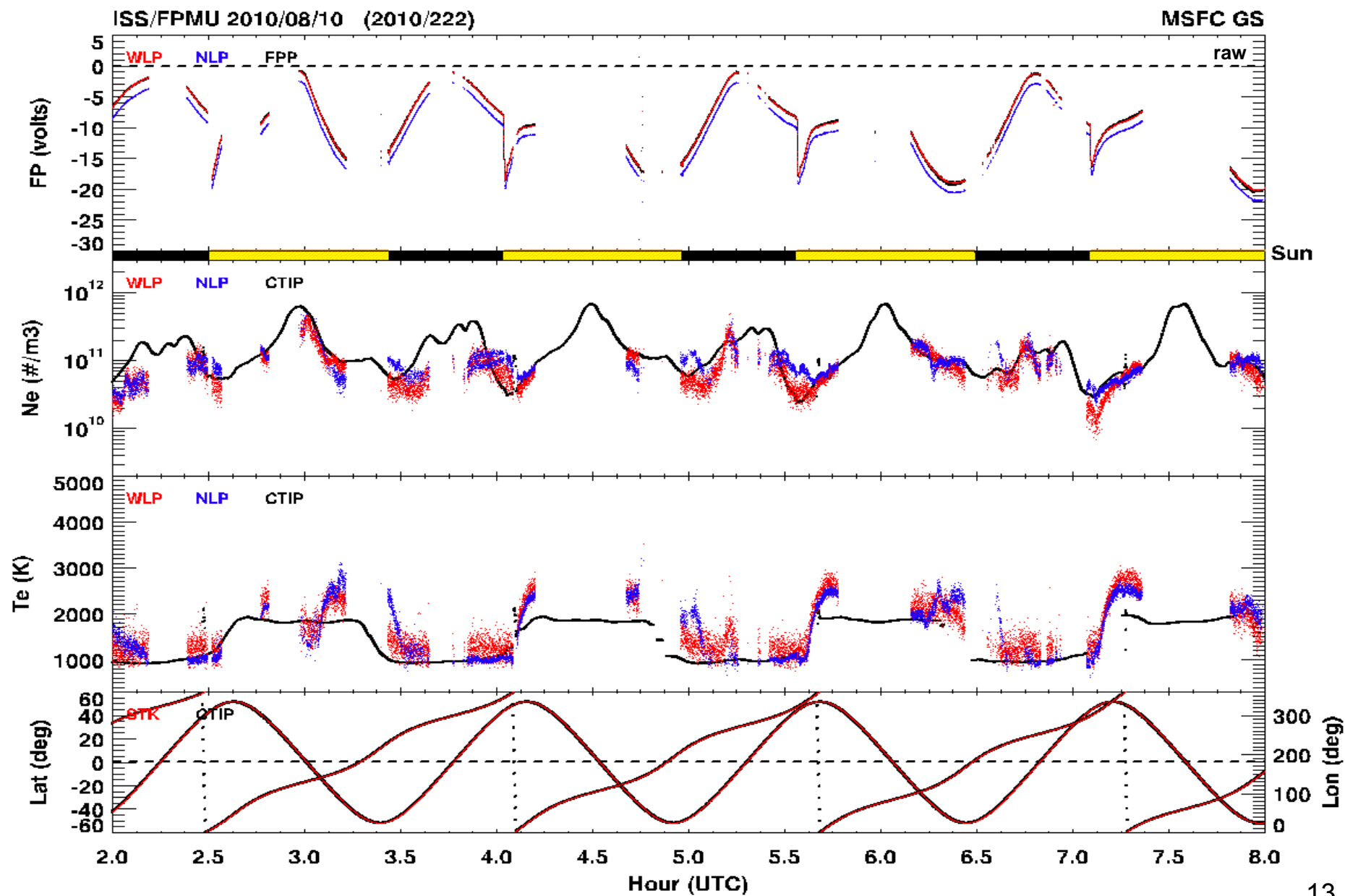


# CTIPe\_RT Output and FPMU Survey Plot



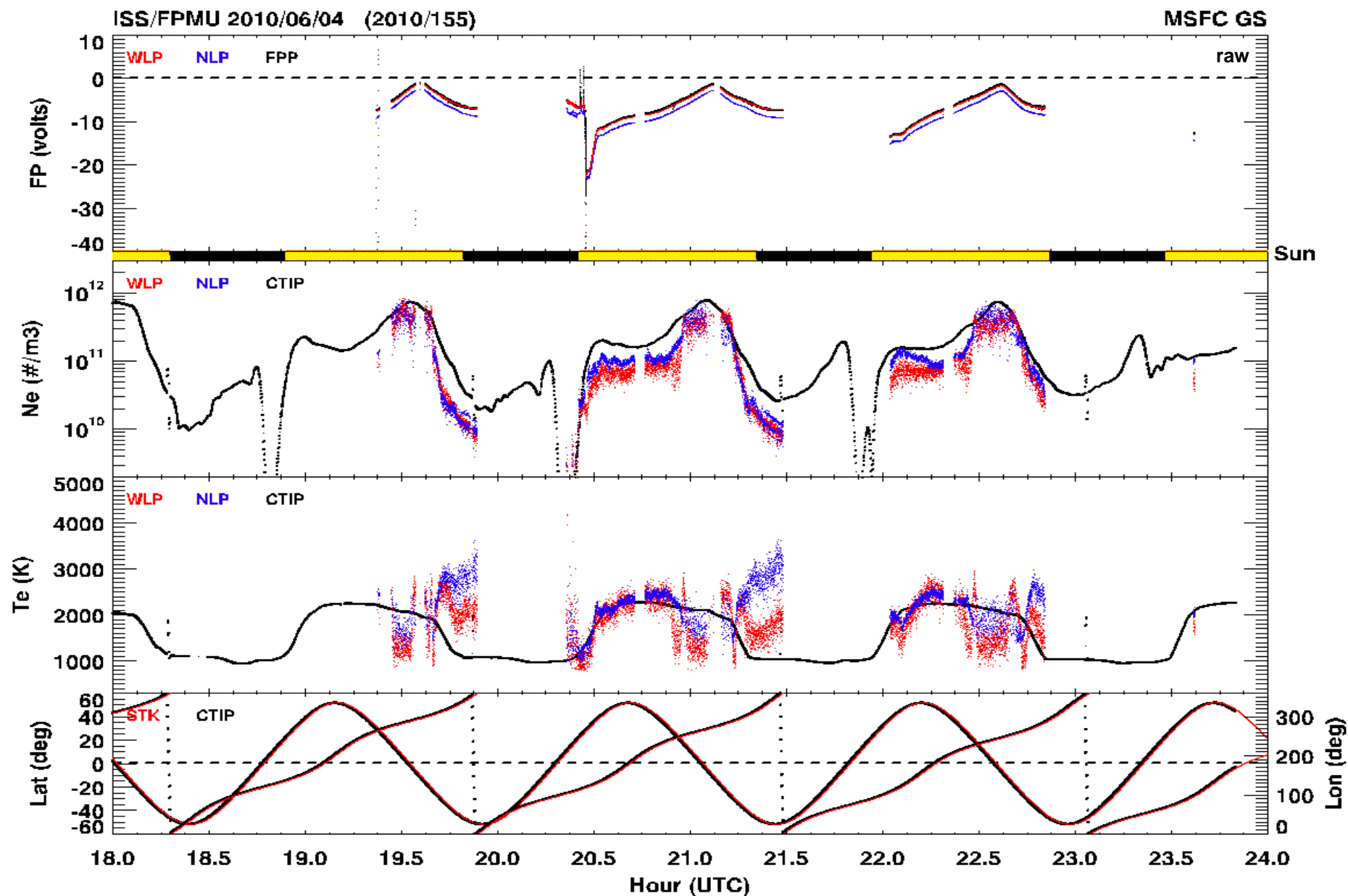


# Good CTIPe\_RT/FPMU Comparison





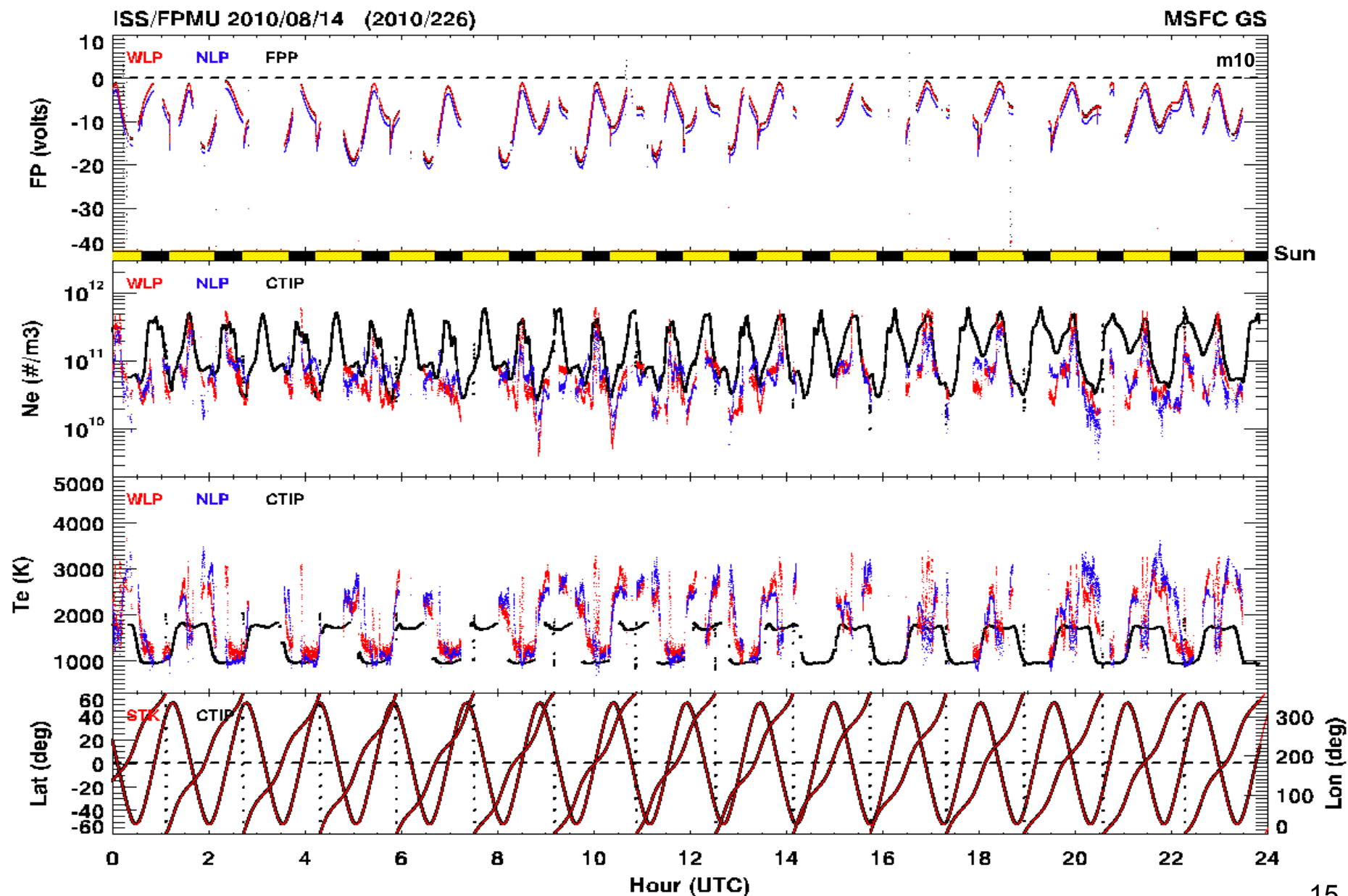
# Good CTIPe\_RT/FPMU Comparison





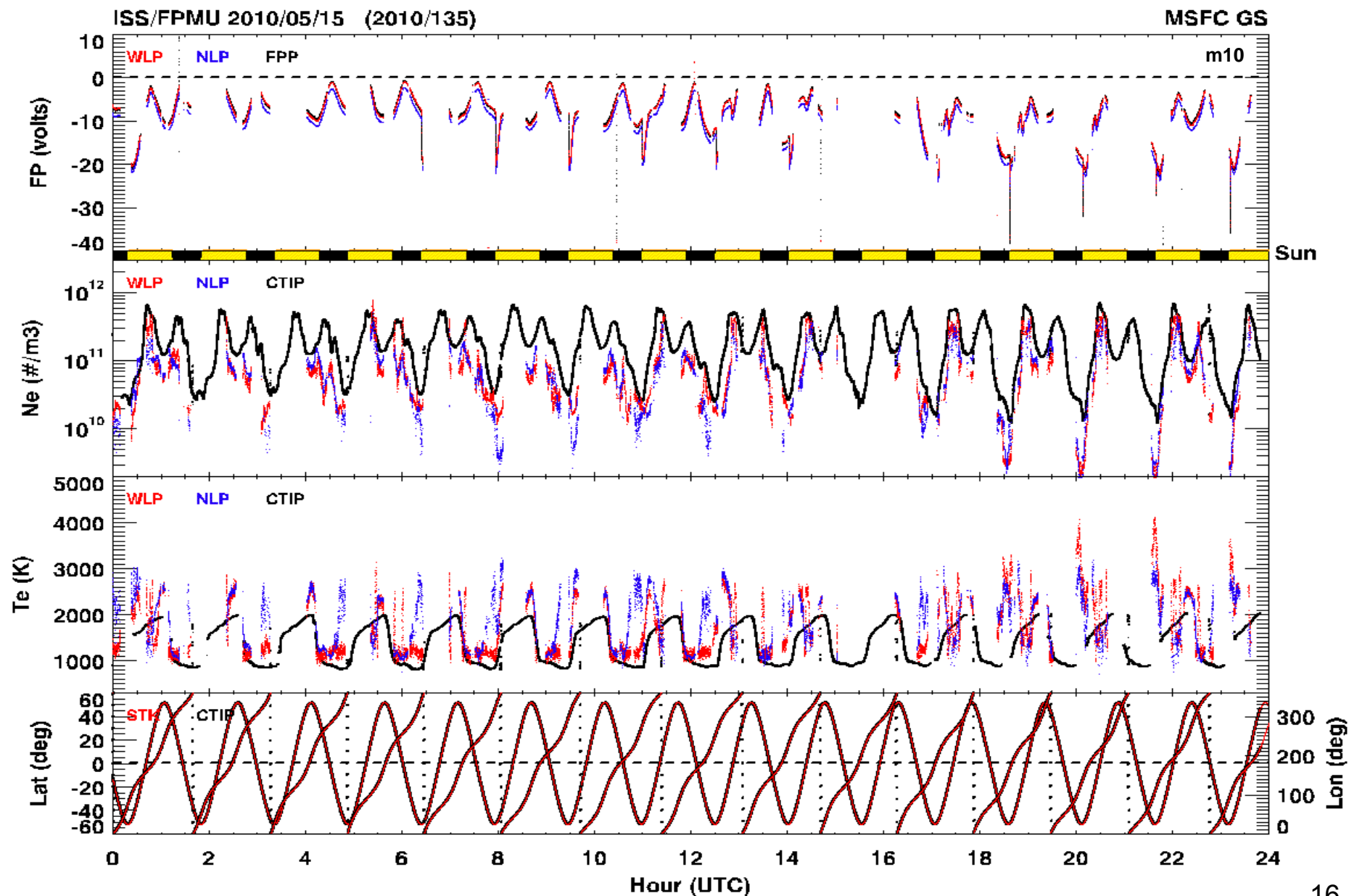


# Less Good CTIPe\_RT/FPMU Comparison





# Poor CTIPe\_RT/FPMU Comparison



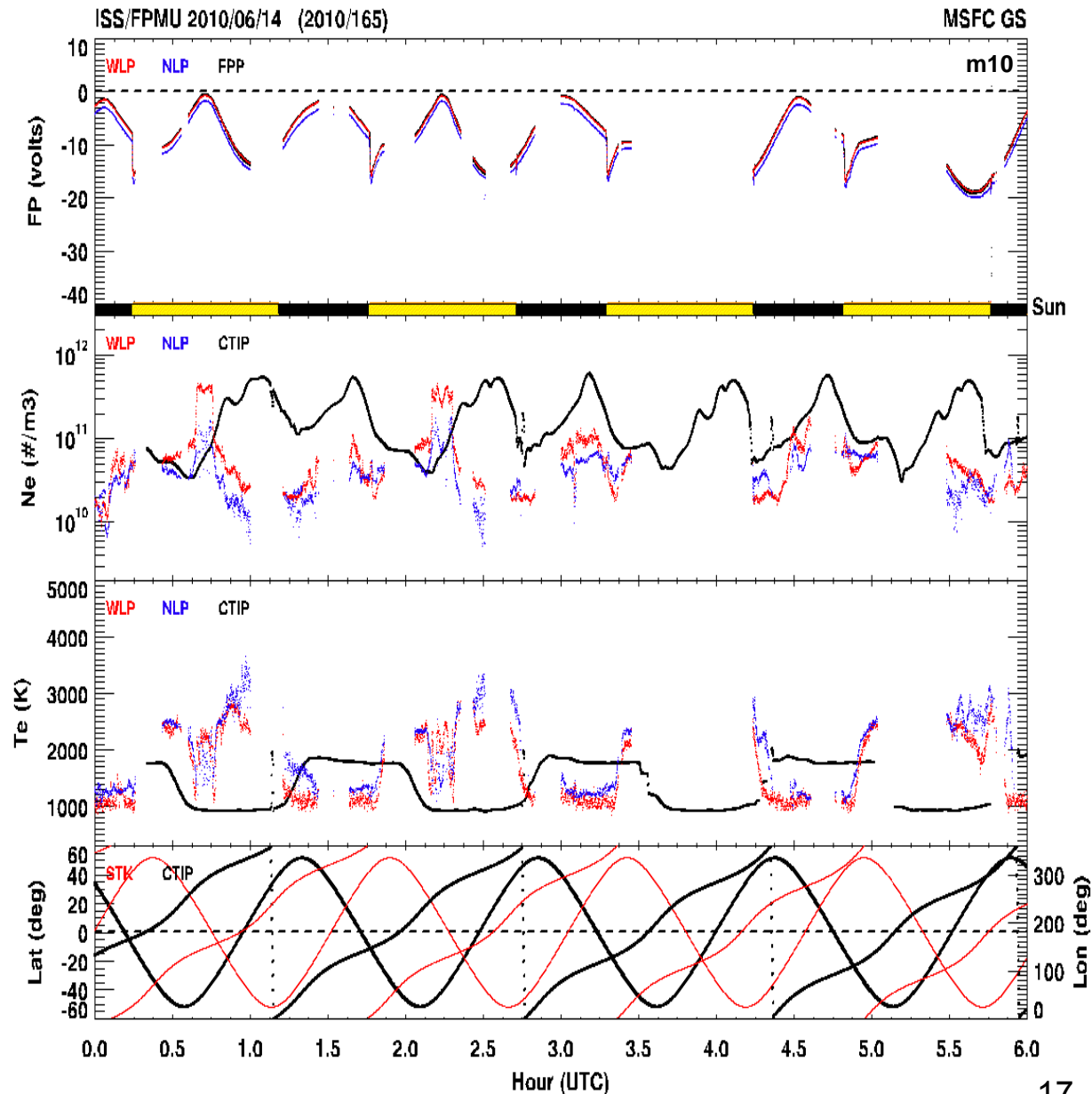
# Incorrect CTIPe\_RT ISS Ephemeris Issue

## MSFC FPMU analysis software

- ISS ephemeris generated from NORAD TLE's using Satellite Tool Kit (STK) software is consistent with ISS/FPMU data

## CCMC real time model

- CTIP output based on incorrect ISS ephemeris obtained from SSCWeb for complete 2010/165 – 168 ISR World Day campaign
- CCMC considering options for robust orbit generation tools for real time model support

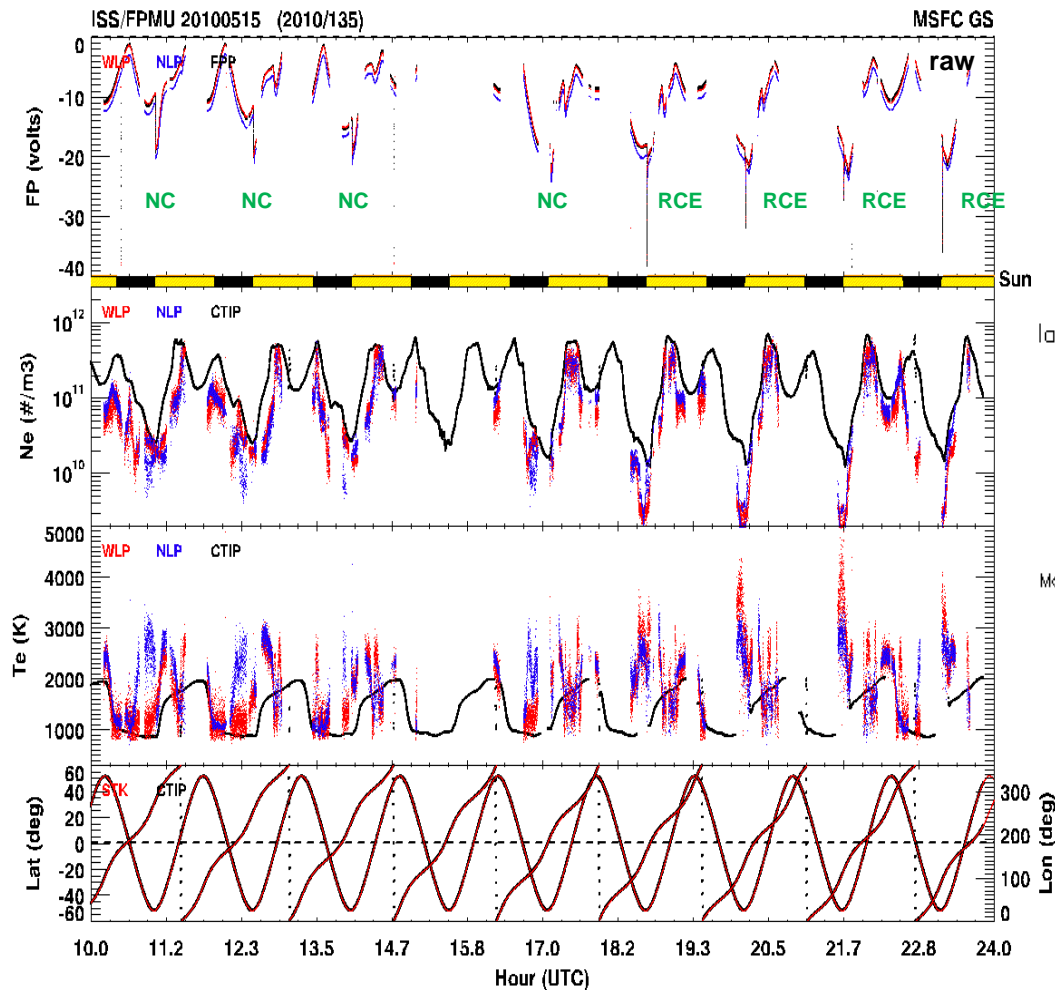




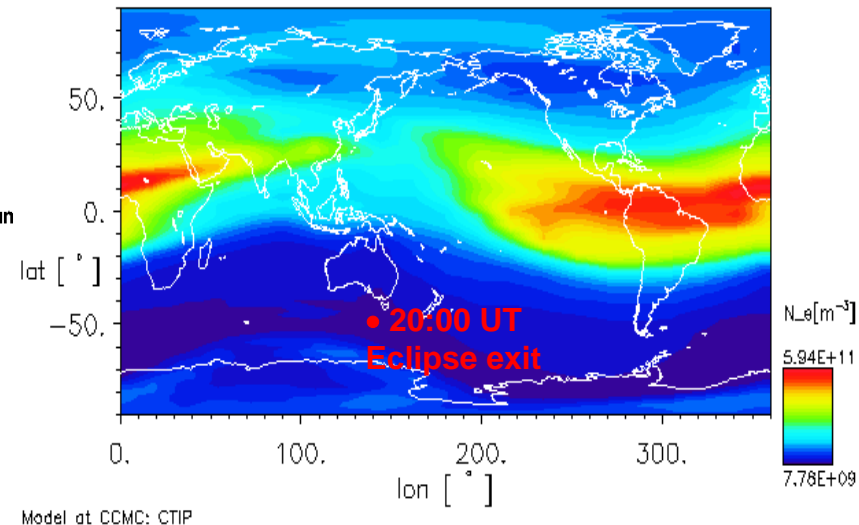


# Characterizing High Latitude Charging Environment

- ISS environments teams are investigating vehicle charging observed at eclipse exit
- CTIPe\_RT model confirmed physical origin of the plasma depletions for charging events observed at high latitudes, allows us to predict periods for studying charging phenomenon



05/15/2010 Time = 20:00:00 UT H= 360.km



**10-17 UT Eclipse Exit**  
Normal charging (NC) events  
observed at eclipse exit

**17 – 24 UT Eclipse Exit**  
Rapid charging (RC) events observed  
when eclipse exit occurs in low  
density plasma troughs



# FPMU Data Unavailable During EVA

